Optics development for the Con-X/SXT telescope at OAB (Italy): a status report

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- review of the activities related to the SXT optics development on course at OAB;
- results obtained with Alumina carriers;
- results obtained with SiC carriers;
- future plans.

Project short overview:

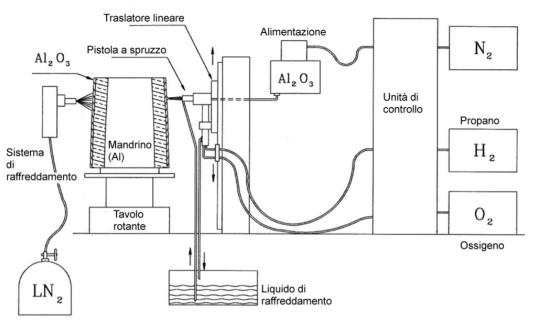
OAB is currently investigating the SXT optics realization option based on the use of carriers made of ceramic materials for making integral mirror shells with Wolter-I profile. Compared to Ni, ceramic materials present the great advantage of a much lower density (a factor > 2.6)

- Mirror shells production method: epoxy replication
- Materials under investigation for making the carriers:
 - CVD Silicon Carbide (SiC):
 (cost at present large but high thermal-mechanical characteristics);
 - plasma-sprayed Alumina (Al₂O₃) (low cost but poorer mechanical parameters).

The Con-X-SXT past and present activities are carried out exploiting the mandrel with polynomial profile fabricated to prove the feasibility of the WFXT mission. $\emptyset = 60 \text{ cm}$ Total Height: 24 cm (to fit the WFXT wide field application)

N.B.: the low mirror height tends to enhance the out-of-phase roundness errors of the shell in determining the optics imaging quality N.B.#2: the profile of the WFXT manufactured mandrel with polynomial shape introduces an intrinsic error for on-axis photons of 8 arcsec HEW.

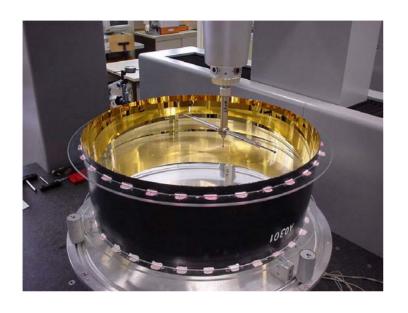
Realization of *Alumina* carriers by Plasma Spray: a low cost process working at room temperature

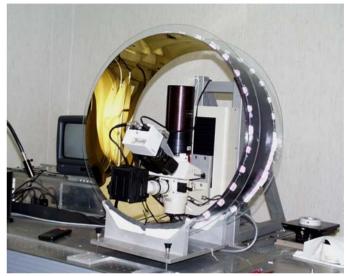






Thin wall Al₂O₃ mirror shell



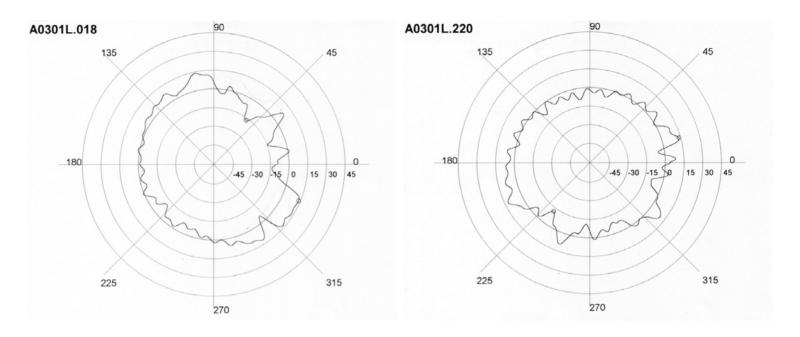


■ Diameter: 600 mm

■ Thickness: 1 mm

■ The thickness to diameter ratio is near to the weight requirements of Con-X/SXT

Thin wall Al₂O₃ shell circularities

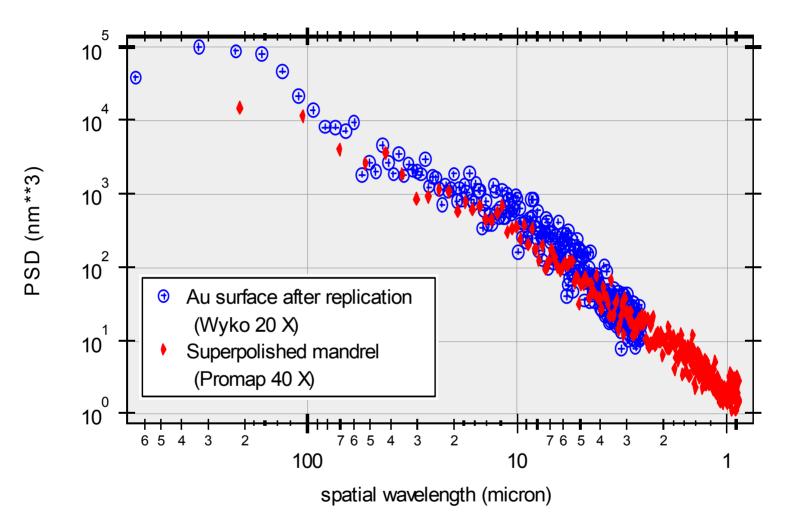


Hyperbole side

Parabola side

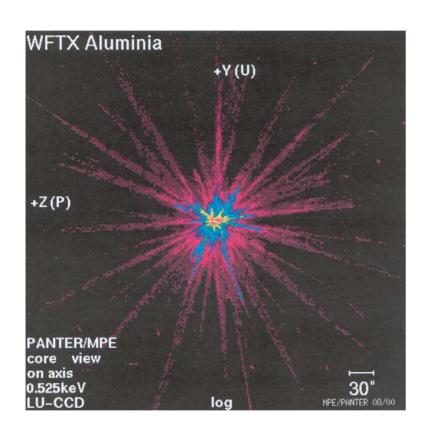
- The shell has been replicated using stiffening glass rings with 32 glued points
- The mean P-V of the waves is about 5-6 microns
- The waviness of the profile match the positions of the glued points
- The next replication will be done using a continuous line of glue

Microroughness: Au mirror shell vs superpolished mandrel



Spatial Wavelength Window: $3 \mu m - 200 \mu m$: σ -mandrel = 2.8 Å σ -mirror = 4.6 Å

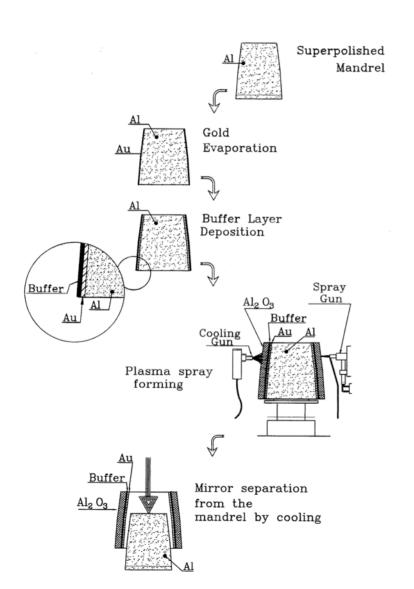
X-Ray imaging tests performed at the Panter MPE facility

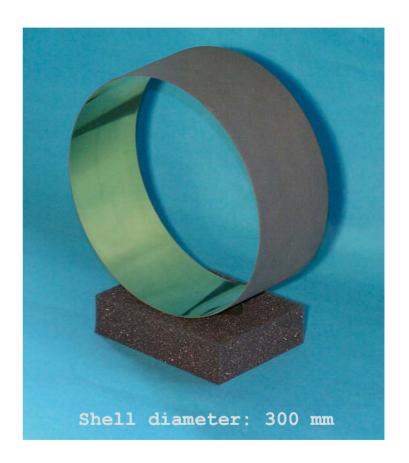


In-focus image at 0.525 KeV (HEW=45 arcsec)

50 mm intrafocus ring at 1.5 KeV

Al2O3 direct deposition process

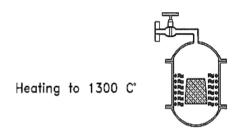


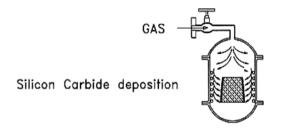


Graphite mandrel

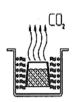


Realization of *SiC* carriers by CVD: <u>a process</u> working at high temperature (at present quite expensive)





Graphite burn out



Grinding



Silicon Carbide (SiC) Carrier

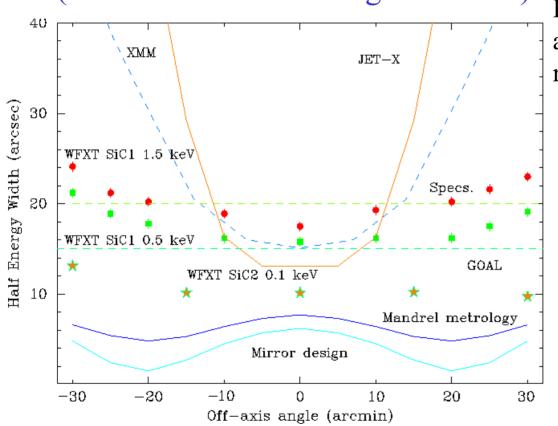


SiC shells produced for thefeasibility study of the WFXT mission concept

•Sic-1: 3 mm thick Weight: 5.4 Kg

• Sic-2: 2 mm thick Weight: 3 Kg



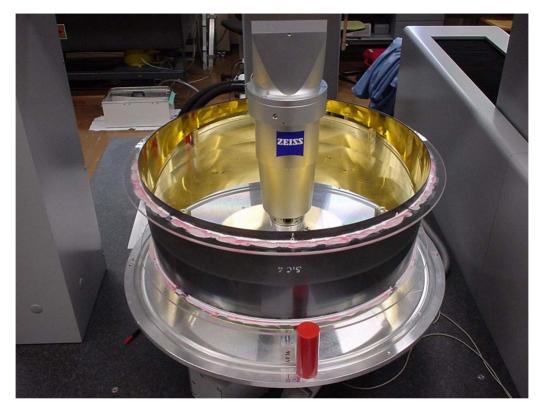


The Sic-2 shell has reached an HEW of 10 arcsec in very good agreement with the SXT requirements

In October 2000 a third shell with a thickness of 1 mm and a weight of 1.4 Kg has been successfully replicated.

The matured experience with these mirror shells can be exploited also for the Constellation-X SXT development.

Thin wall SiC mirror shell (SiC # 4)



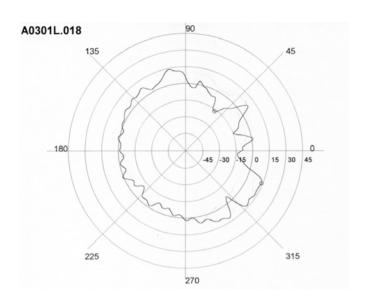
• Diameter: 600 mm;

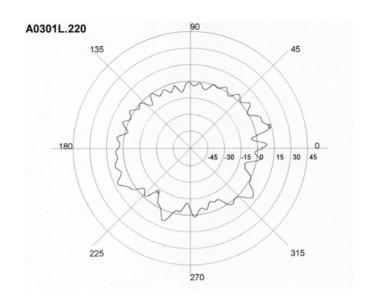
• Thickness: 1 mm;

• Weight: 1.4 *Kg*;

• *Thickness-to-diameter* ratio very close to the Con-X/SXT mass requirements.

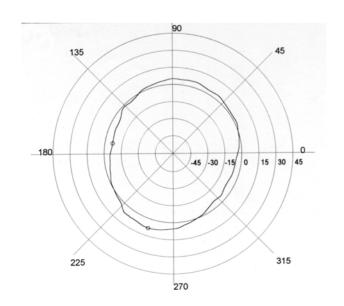
Thin wall (1.0 mm) Al₂O₃ vs. SiC shell circularities



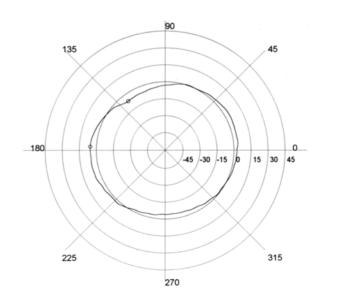


Al₂O₃

Hyperbole side

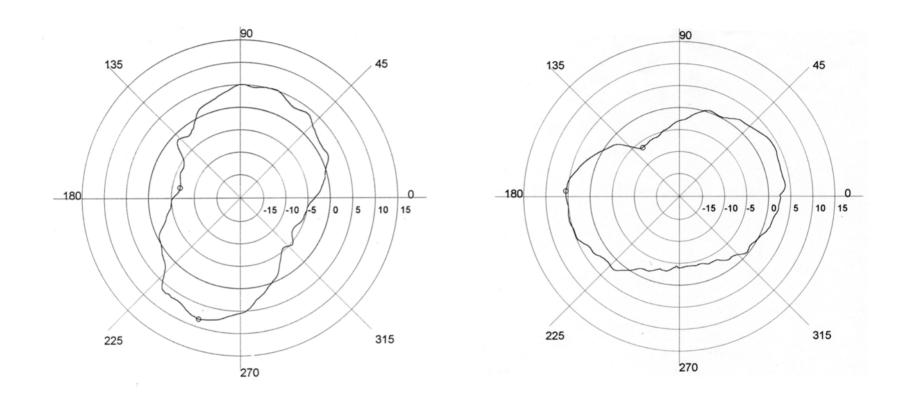


Parabola side



SiC

Thin wall (1 mm) SiC shell circularities *

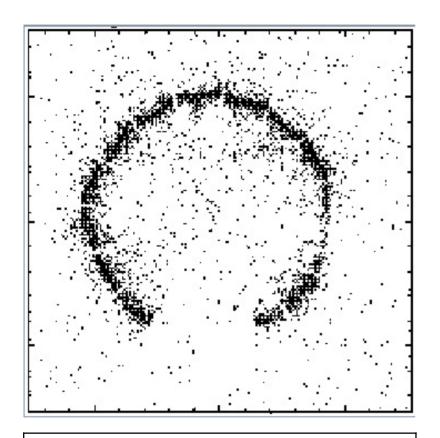


Hyperbole side

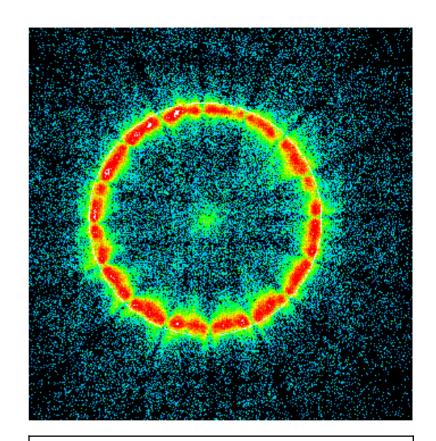
Parabola side

^{*} Scale in microns

Intrafocus rings at 1.5 KeV

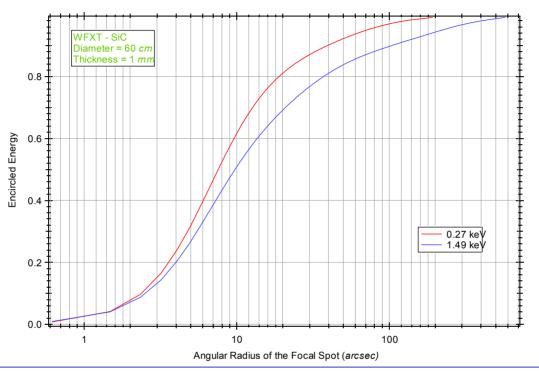


SiC-2 Shell (2 mm thick) Marshall XRCF (1999)

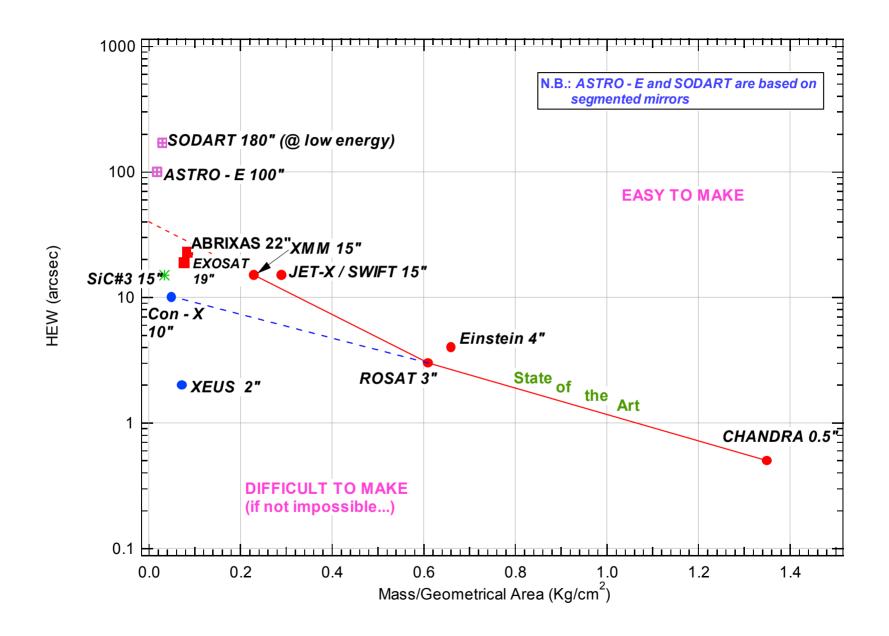


SiC-4 Shell (1 mm thick) Panter MPE (2001)

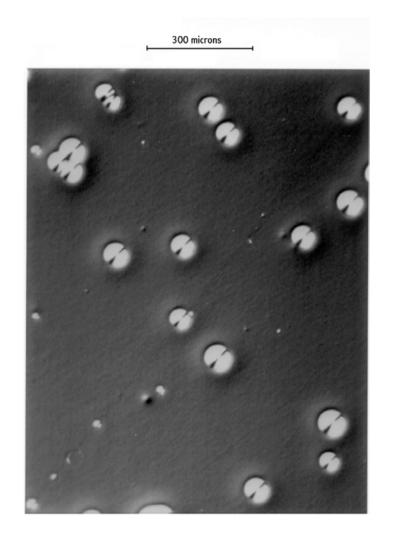
Encircled Energy Function and HEW derived by the X-Ray imaging tests performed at the Panter-MPE facility (March 2001)

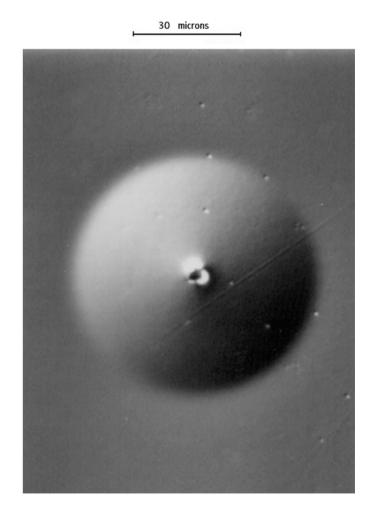


Energy (keV)	1 <i>mm</i> shell - Panter 2001 HEW (<i>arcsec</i>)	2 <i>mm</i> shell - Marshall 1999 HEW (<i>arcsec</i>)
0.108		10.5 (aver.)
0.277	<i>15.12</i>	13.5 (aver.)
1.49	19.97	24.9 (aver.)



Nomarsky microscope images showing the presence of microbubbles on the MS gold surface (perhaps probably due to water vapour contamination during the mirror shell epoxy replication)





Possible Improvements:

- ↓ elimination of the presence of microbubbles on the mirror Au surface avoiding the contamination by H₂O vapor during the epoxy replication
 ↓ the process will be performed into a more "clean" environment;
- With a Con-X mandrel the error contribution will be reduced by a factor 2.5 (assuming the same circularity deviation);
- ↓ use of mandrels with a more favorable geometrical figure

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intrinsic on-axis HEW error introduced by the mandrel

WFXT mandrel

Zeiss - Con-X 50 cm mandrels:

8 arcsec

VS

5 arcsec
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Future Plans

- replication of a second SiC shell 1 mm thick exploiting the WFXT 60 cm diam. mandrel in order to verify:
 - * the repeatability and possible improvements of the results already achieved;
 - the imaging performances after having removed the two stiffening rings.
- realization of a mirror shell exploiting one of the two 50 cm diam.
 mandrels manufactured by Zeiss for the Con-X / SXT feasibility study.

An investigation aimed to verify the possibility of using the reaction bonding technique to produce the SiC carries (a much cheaper approach compared to the CVD method utilized until now) is presently on course.